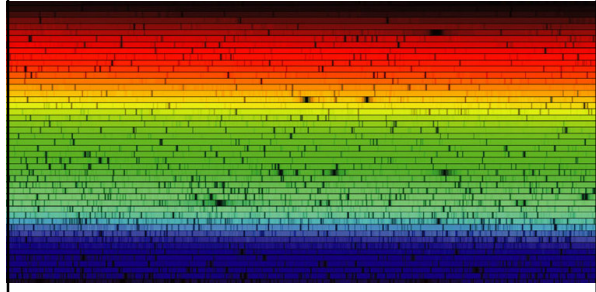


## Chapter 5 Light and Matter: Reading Messages from the Cosmos



## Agenda

- Announce:
  - Light & Spectroscopy Tutorial Due by next test (counted as HW)
  - Thin Lens Lab on Thursday
- Projects
- Ch. 5—Light & Matter
- Elegant Universe

## 5.1 Light in Everyday Life

- How do we experience light?
- How do light and matter interact?

## How do we experience light?

- The warmth of sunlight tells us that light is a form of energy
- We can measure the flow of energy (power) in light in units of **watts**: 1 watt = 1 joule/s

## What do you pay for?

- Watts
- Joules

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## Colors of Light

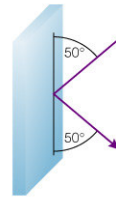


- White light is made up of many different colors

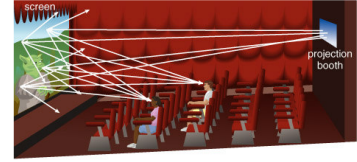
## How do light and matter interact?

- Emission
- Absorption
- Transmission
  - Transparent objects transmit light
  - Opaque objects block (absorb) light
- Reflection or Scattering

## Reflection and Scattering

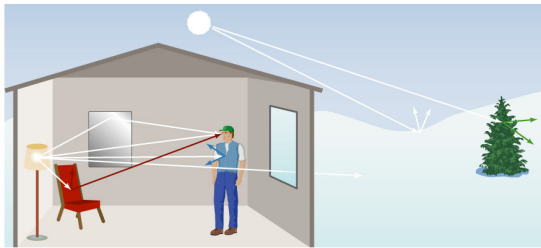


Mirror reflects light in a particular direction



Movie screen scatters light in all directions

## Interactions of Light with Matter



Interactions between light and matter determine the appearance of everything around us

## Thought Question Why is a rose red?

- The rose absorbs red light.
- The rose transmits red light.
- The rose emits red light.
- The rose reflects red light.

## Thought Question Why is a rose red?

- The rose absorbs red light.
- The rose transmits red light.
- The rose emits red light.
- d) The rose reflects red light.**

## What have we learned?

- How do we experience light?
  - Light is a form of energy
  - Light comes in many colors that combine to form white light.
- How does light interact with matter?
  - Matter can emit light, absorb light, transmit light, and reflect (or scatter) light.
  - Interactions between light and matter determine the appearance of everything we see.

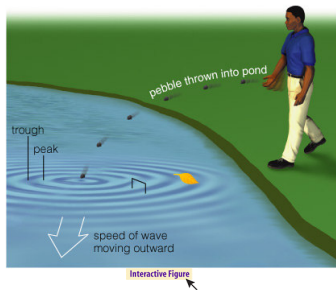
## 5.2 Properties of Light

- What is light?
- What is the electromagnetic spectrum?

## What is light?

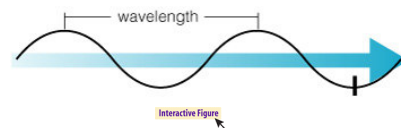
- Light can act either like a wave or like a particle
- Particles of light are called **photons**

## Waves



- A **wave** is a pattern of motion that can carry energy without carrying matter along with it

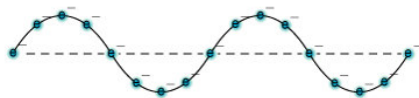
## Properties of Waves



- **Wavelength** is the distance between two wave peaks
- **Frequency** is the number of times per second that a wave vibrates up and down

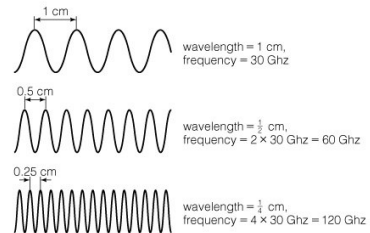
$$\text{wave speed} = \text{wavelength} \times \text{frequency}$$

## Light: Electromagnetic Waves



- A light wave is a vibration of electric and magnetic fields
- Light interacts with charged particles through these electric and magnetic fields

## Wavelength and Frequency



$$\text{wavelength} \times \text{frequency} = \text{speed of light} = \text{constant}$$

## Particles of Light

- Particles of light are called **photons**
- Each photon has a wavelength and a frequency
- The energy of a photon depends on its frequency

## Wavelength, Frequency, and Energy

$$\lambda \times f = c$$

$\lambda$  = wavelength ,  $f$  = frequency

$c = 3.00 \times 10^8$  m/s = speed of light

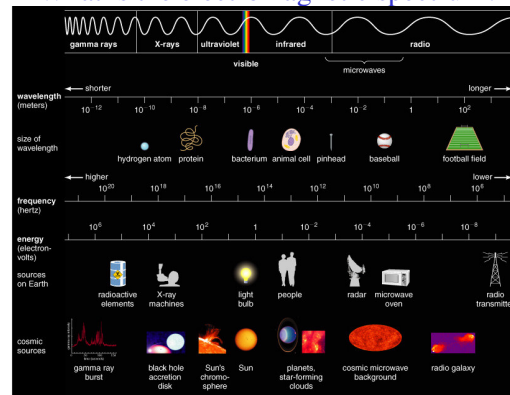
$E = h \times f$  = photon energy

$h = 6.626 \times 10^{-34}$  joule x s

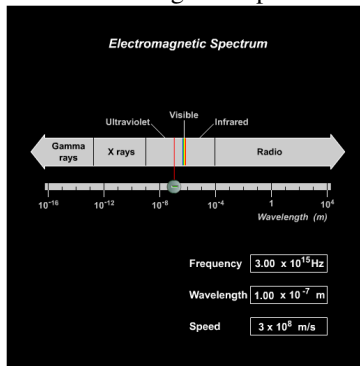
## Special Topic: Polarized Sunglasses

- **Polarization** describes the direction in which a light wave is vibrating
- Reflection can change the polarization of light
- Polarized sunglasses block light that reflects off horizontal surfaces

## What is the electromagnetic spectrum?



## The Electromagnetic Spectrum



## Thought Question

The higher the photon energy...

- the longer its wavelength.
- the shorter its wavelength.
- energy is independent of wavelength.

## Thought Question

### The higher the photon energy...

- a) the longer its wavelength.
- b) the shorter its wavelength.**
- c) energy is independent of wavelength.

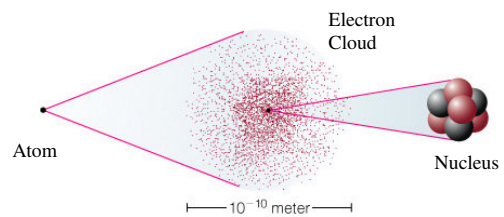
## What have we learned?

- **What is light?**
  - Light can behave like either a wave or a particle
  - A light wave is a vibration of electric and magnetic fields
  - Light waves have a wavelength and a frequency
  - Photons are particles of light.
- **What is the electromagnetic spectrum?**
  - Human eyes cannot see most forms of light.
  - The entire range of wavelengths of light is known as the electromagnetic spectrum.

## 5.3 Properties of Matter


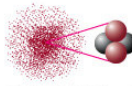

- What is the structure of matter?
- What are the phases of matter
- How is energy stored in atoms?

## What is the structure of matter?



## Atomic Terminology




- **Atomic Number** = # of protons in nucleus
- **Atomic Mass Number** = # of protons + neutrons

Hydrogen ( ${}^1\text{H}$ )	Helium ( ${}^4\text{He}$ )	Carbon ( ${}^{12}\text{C}$ )
		
atomic number = 1 atomic mass number = 1 (1 electron)	atomic number = 2 atomic mass number = 4 (2 electrons)	atomic number = 6 atomic mass number = 12 (6 electrons)

- **Molecules:** consist of two or more atoms ( $\text{H}_2\text{O}$ ,  $\text{CO}_2$ )

## Atomic Terminology

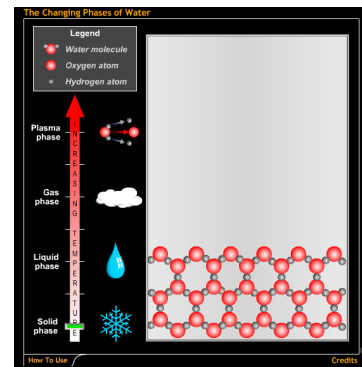
- **Isotope:** same # of protons but different # of neutrons. ( ${}^4\text{He}$ ,  ${}^3\text{He}$ )

Isotopes of Carbon		
carbon-12	carbon-13	carbon-14
		
${}^{12}\text{C}$ (6 protons + 6 neutrons)	${}^{13}\text{C}$ (6 protons + 7 neutrons)	(6 protons + 8 neutrons)

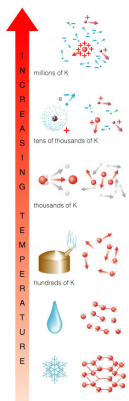
## What are the phases of matter?

- Familiar phases:
  - Solid (ice)
  - Liquid (water)
  - Gas (water vapor)
- Phases of same material behave differently because of differences in chemical bonds

## Phases of Water

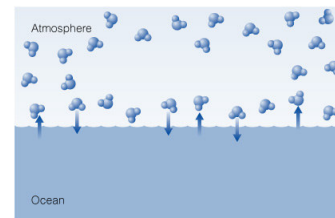


## Phase Changes



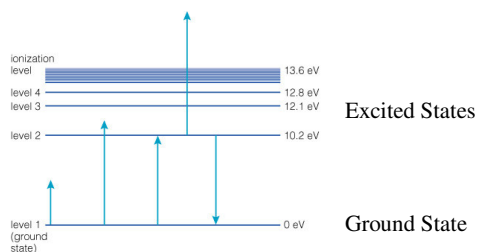
- **Ionization:** Stripping of electrons, changing atoms into plasma
- **Dissociation:** Breaking of molecules into atoms
- **Evaporation:** Breaking of flexible chemical bonds, changing liquid into solid
- **Melting:** Breaking of rigid chemical bonds, changing solid into liquid

## Phases and Pressure



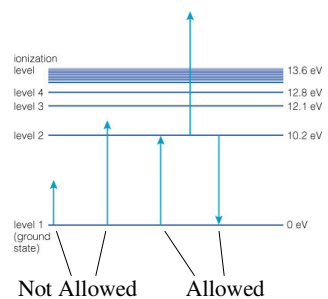
- Phase of a substance depends on both temperature and pressure
- Often more than one phase is present

## How is energy stored in atoms?



- Electrons in atoms are restricted to particular energy levels

## Energy Level Transitions



- The only allowed changes in energy are those corresponding to a transition between energy levels

## What have we learned?

- What is the structure of matter?
  - Matter is made of atoms, which consist of a nucleus of protons and neutrons surrounded by a cloud of electrons
- What are the phases of matter?
  - Adding heat to a substance changes its phase by breaking chemical bonds.
  - As temperature rises, a substance transforms from a solid to a liquid to a gas, then the molecules can dissociate into atoms
  - Stripping of electrons from atoms (ionization) turns the substance into a plasma

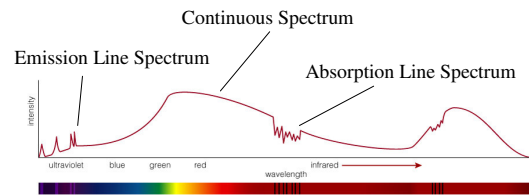
## What have we learned?

- How is energy stored in atoms?
  - The energies of electrons in atoms correspond to particular energy levels.
  - Atoms gain and lose energy only in amount corresponding to particular changes in energy levels.

## 5.4 Learning from Light

- What are the three basic types of spectra?
- How does light tell us what things are made of?
- How does light tell us the temperatures of planets and stars?
- How do we interpret an actual spectrum?

## What are the three basic types of spectra?



Spectra of astrophysical objects are usually combinations of these three basic types

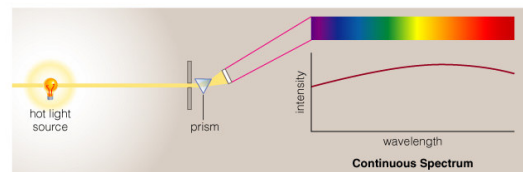
## Three Types of Spectra

**The Details of Spectra - Illustrating Kirchhoff's Laws**

<p><b>Show</b> Continuous Spectrum</p> <p>The light bulb produces light with a continuous spectrum.</p>	<p>The spectrum shows a smooth, continuous rainbow of light.</p> <p>A graph of the spectrum is also continuous, notice the intensity varies slightly at different wavelengths.</p>
<p><b>Show</b> Emission Line Spectrum</p> <p>The cloud also emits its own light, but only at specific wavelengths determined by its composition.</p>	<p>We see bright emission lines at specific wavelengths (color), but no other light.</p> <p>The graph shows an upward spike at the wavelength of each emission line.</p>
<p><b>Show</b> Absorption Line Spectrum</p> <p>The cloud absorbs light at specific wavelengths determined by its composition.</p>	<p>We see dark absorption lines where the cloud has absorbed lights of specific wavelengths (colors).</p> <p>The graph shows a dip in intensity at the wavelength of each absorption line.</p>

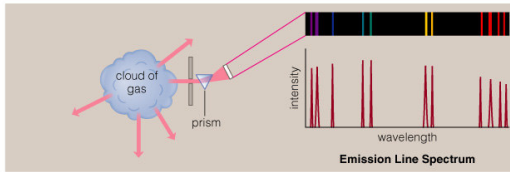
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## Continuous Spectrum



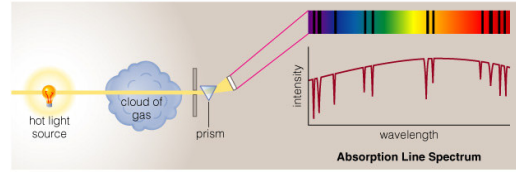
- The spectrum of a common (incandescent) light bulb spans all visible wavelengths, without interruption

## Emission Line Spectrum



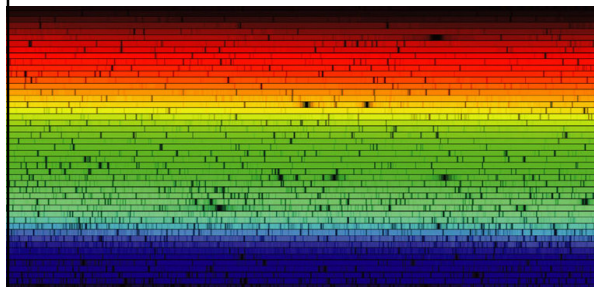
- A thin or low-density cloud of gas emits light only at specific wavelengths that depend on its composition and temperature, producing a spectrum with bright emission lines

## Absorption Line Spectrum



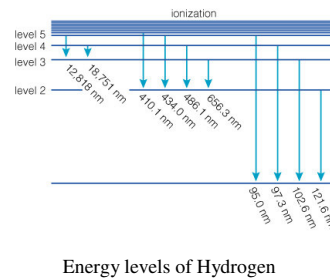
- A cloud of gas between us and a light bulb can absorb light of specific wavelengths, leaving dark absorption lines in the spectrum

## How does light tell us what things are made of?



Spectrum of the Sun

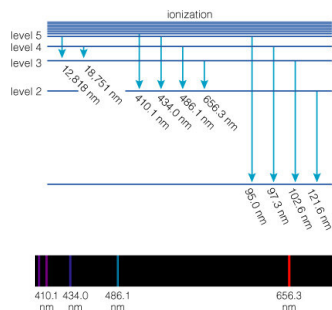
## Chemical Fingerprints



- Each type of atom has a unique set of energy levels
- Each transition corresponds to a unique photon energy, frequency, and wavelength

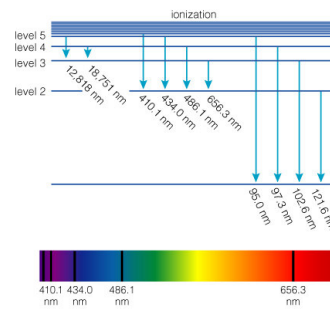
Energy levels of Hydrogen

## Chemical Fingerprints



- Downward transitions produce a unique pattern of emission lines

## Chemical Fingerprints



- Because those atoms can absorb photons with those same energies, upward transitions produce a pattern of absorption lines at the same wavelengths

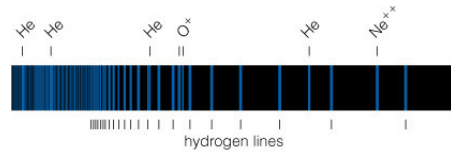


## Chemical Fingerprints



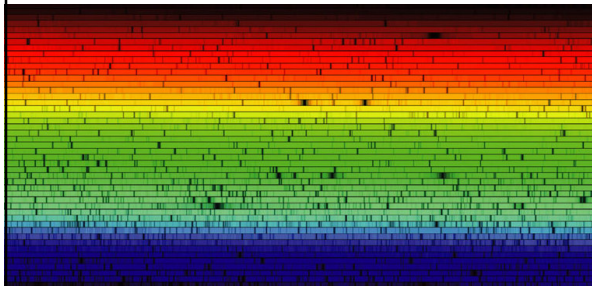
- Each type of atom has a unique spectral fingerprint

## Chemical Fingerprints

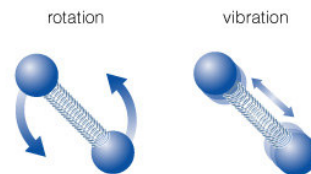


- Observing the fingerprints in a spectrum tells us which kinds of atoms are present

## Example: Solar Spectrum



## Energy Levels of Molecules



- Molecules have additional energy levels because they can vibrate and rotate

## Energy Levels of Molecules

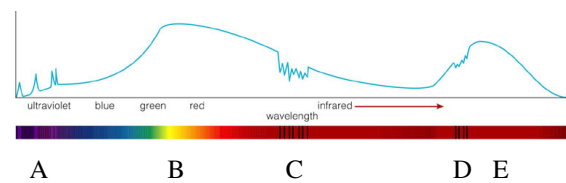


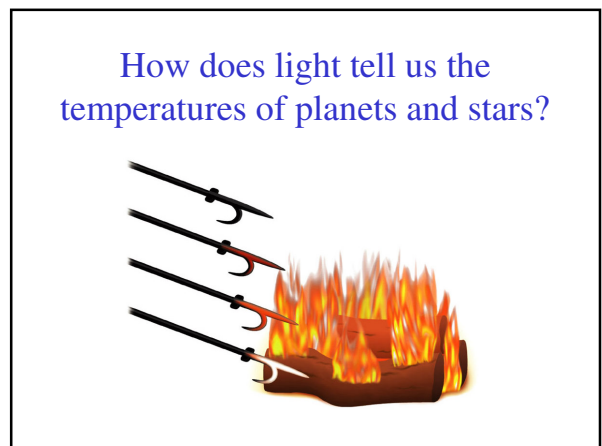
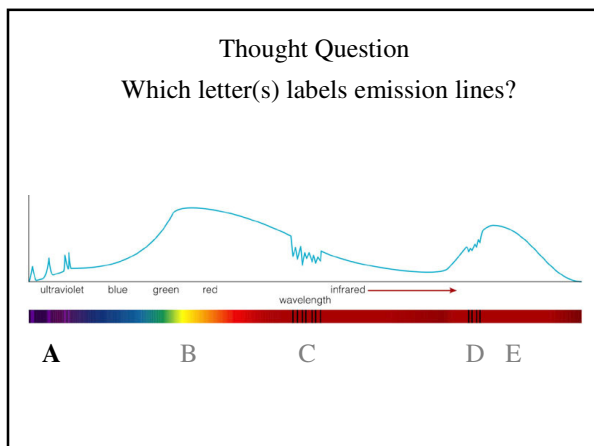
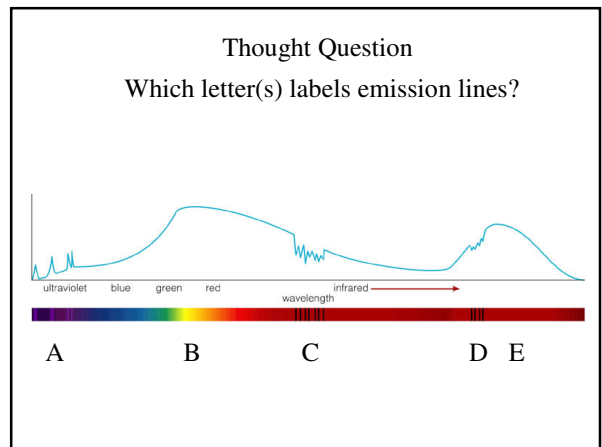
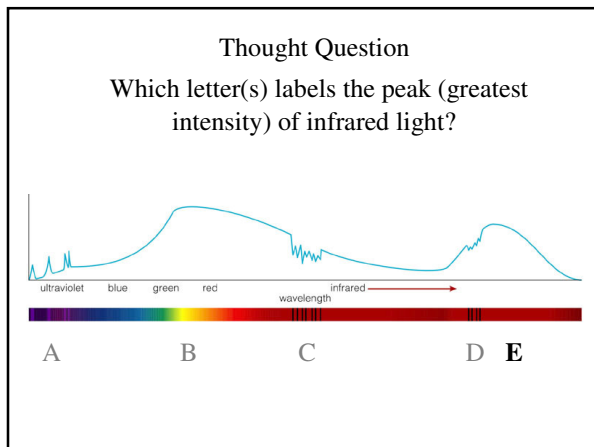
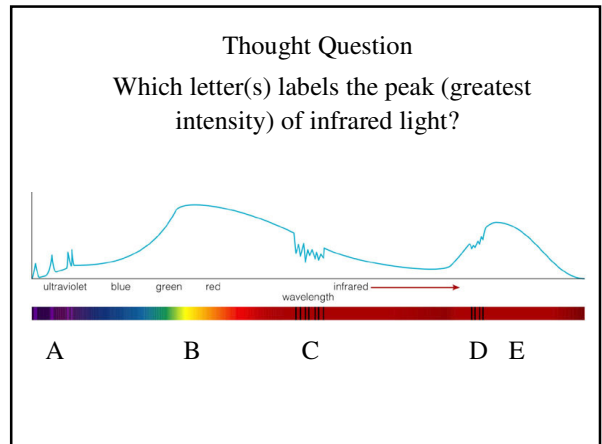
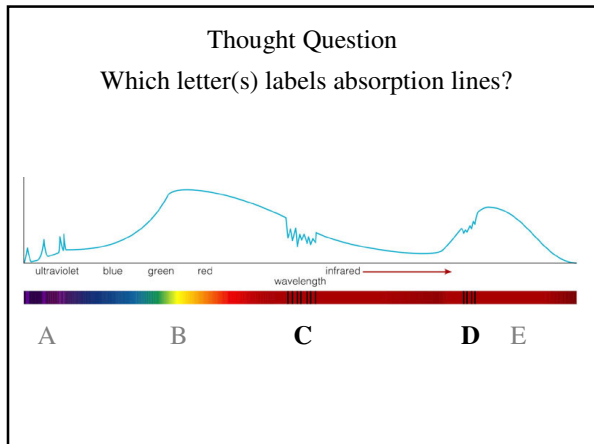
Spectrum of Molecular Hydrogen

- The large numbers of vibrational and rotational energy levels can make the spectra of molecules very complicated
- Many of these molecular transitions are in the infrared part of the spectrum

## Thought Question

Which letter(s) labels absorption lines?



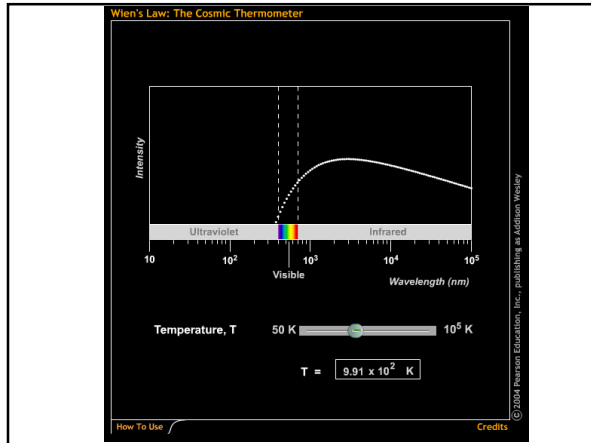
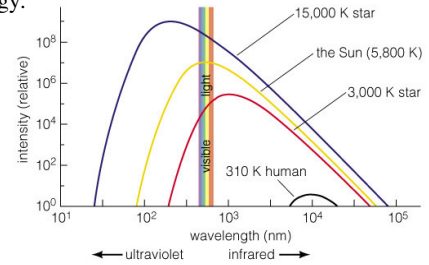


## Thermal Radiation

- Nearly all large or dense objects emit thermal radiation, including stars, planets, you...
- An object's thermal radiation spectrum depends on only one property: its **temperature**

## Properties of Thermal Radiation

- Hotter objects emit more light at all frequencies per unit area.
- Hotter objects emit photons with a higher average energy.



## Thought Question

### Which is hotter?

- A blue star.
- A red star.
- A planet that emits only infrared light.

## Thought Question

### Which is hotter?

- A blue star.
- A red star.
- A planet that emits only infrared light.

## Thought Question

### Why don't we glow in the dark?

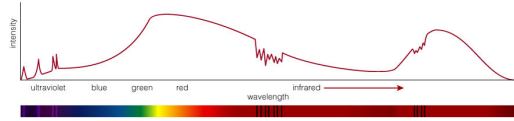
- People do not emit any kind of light.
- People only emit light that is invisible to our eyes.
- People are too small to emit enough light for us to see.
- People do not contain enough radioactive material.

## Thought Question

### Why don't we glow in the dark?

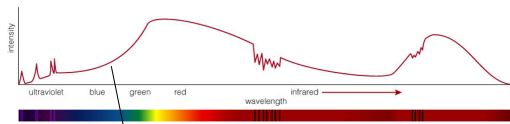
- People do not emit any kind of light.
- People only emit light that is invisible to our eyes.**
- People are too small to emit enough light for us to see.
- People do not contain enough radioactive material.

## How do we interpret an actual spectrum?



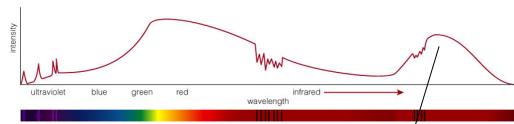
- By carefully studying the features in a spectrum, we can learn a great deal about the object that created it.

## What is this object?



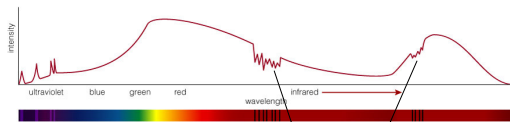
Reflected Sunlight:  
Continuous spectrum of visible light is like the Sun's except that some of the blue light has been absorbed - object must look red

## What is this object?



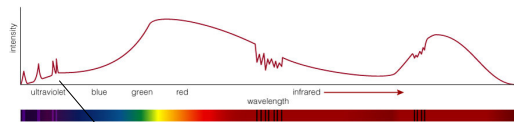
Thermal Radiation:  
Infrared spectrum peaks at a wavelength corresponding to a temperature of 225 K

## What is this object?



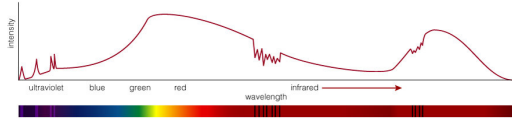
Carbon Dioxide:  
Absorption lines are the fingerprint of  $\text{CO}_2$  in the atmosphere

## What is this object?



Ultraviolet Emission Lines:  
Indicate a hot upper atmosphere

## What is this object?



Mars!

## What have we learned?

- What are the three basic type of spectra?
  - Continuous spectrum, emission line spectrum, absorption line spectrum
- How does light tell us what things are made of?
  - Each atom has a unique fingerprint.
  - We can determine which atoms something is made of by looking for their fingerprints in the spectrum.

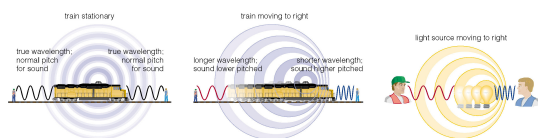
## What have we learned?

- How does light tell us the temperatures of planets and stars?
  - Nearly all large or dense objects emit a continuous spectrum that depends on temperature.
  - The spectrum of that thermal radiation tells us the object's temperature.
- How do we interpret an actual spectrum?
  - By carefully studying the features in a spectrum, we can learn a great deal about the object that created it.

## 5.5 The Doppler Effect

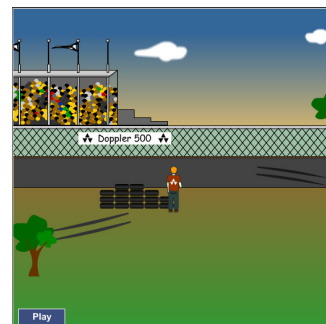
- Our goals for learning
- How does light tell us the speed of a distant object?
- How does light tell us the rotation rate of an object?

## How does light tell us the speed of a distant object?

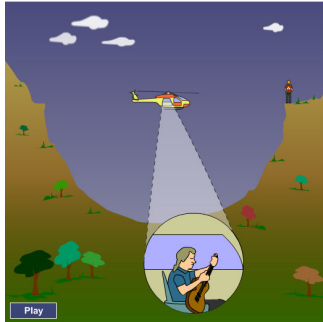


The Doppler Effect

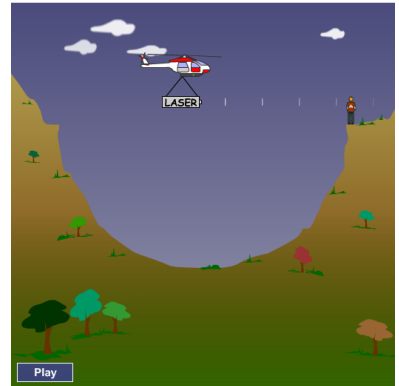
## The Doppler Effect



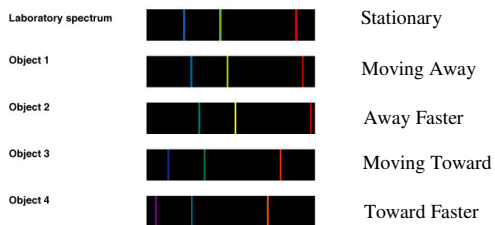
## Explaining the Doppler Effect



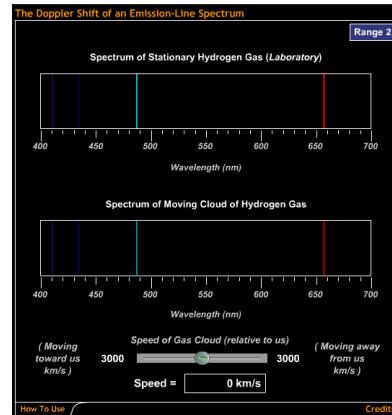
Same for  
Light



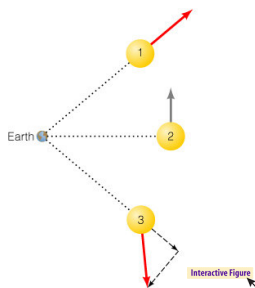
## Measuring the Shift



- We generally measure the Doppler Effect from shifts in the wavelengths of spectral lines



Doppler shift tells us ONLY about the part of an object's motion toward or away from us:



## Thought Question

I measure a line in the lab at 500.7 nm.  
The same line in a star has wavelength 502.8 nm.

What can I say about this star?

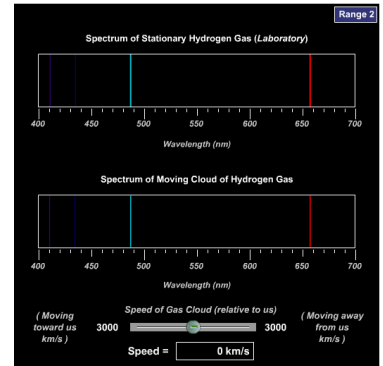
- It is moving away from me.
- It is moving toward me.
- It has unusually long spectral lines.

### Thought Question

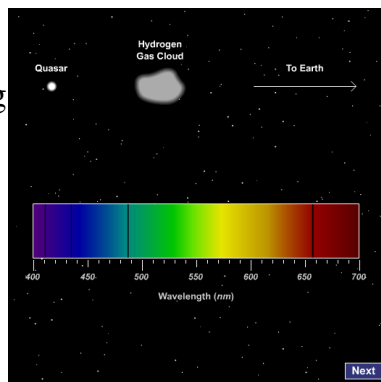
I measure a line in the lab at 500.7 nm.  
The same line in a star has wavelength 502.8 nm.  
What can I say about this star?

- a) It is moving away from me.
- b) It is moving toward me.
- c) It has unusually long spectral lines.

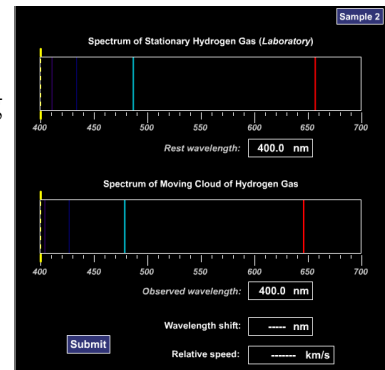
## Measuring Redshift



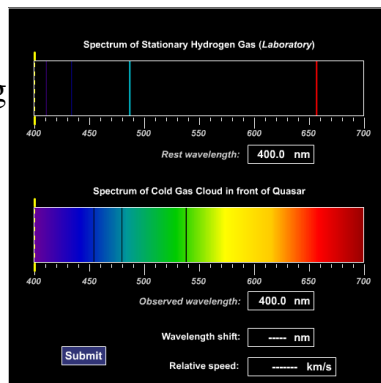
## Measuring Redshift



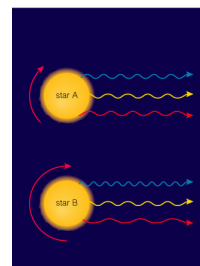
## Measuring Velocity



## Measuring Velocity

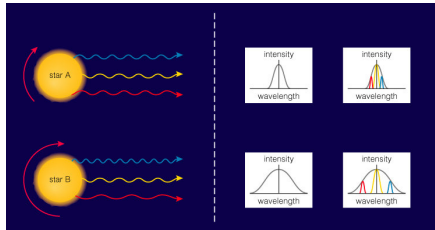


## How does light tell us the rotation rate of an object?



- Different Doppler shifts from different sides of a rotating object spread out its spectral lines

## Spectrum of a Rotating Object



- Spectral lines are wider when an object rotates faster

## What have we learned?

- How does light tell us the speed of a distant object?
  - The Doppler effect tells us how fast an object is moving toward or away from us.
    - **Blueshift**: objects moving toward us
    - **Redshift**: objects moving away from us
- How does light tell us the rotation rate of an object?
  - The width of an object's spectral lines can tell us how fast it is rotating