

The Celestial Sphere • Vast distances to stars prevent us from arrangement Naked eye

- sensing their true 3-D
- observations treat all stars at the same distance, on a giant celestial sphere with the Earth at its center









Diurnal Motion

- Daily motion can be explained by the rotation of the celestial sphere about the north and south celestial poles located directly above the Earth's north and south poles
- The *celestial equator*, which lies directly above the Earth's equator, provides another astronomical reference marker





Annual Motion

- A given star rises 3 minutes 56 seconds earlier each night
- This annual motion is caused by the Earth's motion around the Sun, the result of projection
- The ancients used the periodic annual motion to mark the seasons



The Ecliptic



- The path of the Sun through the stars on the celestial sphere is called the *ecliptic*
- The ecliptic is a projection of the Earth's orbit onto the celestial sphere and is tipped relative to the celestial equator

A • The Earth is closest to the Sun in January, which is unitaring the marking bars

- winter in the northern hemisphereTherefore, the seasons cannot be caused by the
- Sun's proximity to the Earth
 The Earth's *rotation axis* is tilted 23.5° from a line perpendicular to the Earth's orbital plane

















Moon goes through a phases: new, waxing crescent, first quarter, waxing gibbous, full, waning gibbous, third

The Phases of the Moon - The phase cycle is the origin of the month (derived from the word moon) as a time period - The phases of the Moon are caused by the relative positions of the Sun, Earth, and Moon









between the Sun and Moon, with the Earth casting its shadow on the Moon giving it a dull red color







• The tipped orbit allows the shadow of the Earth (Moon) to miss the Moon (Earth)

