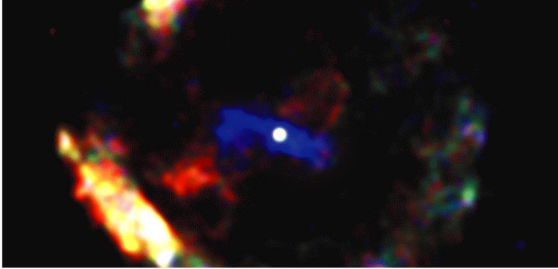


Chapter 18 The Bizarre Stellar Graveyard



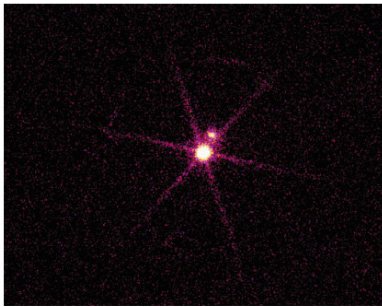
© 2004 Pearson Education, Inc., publishing as Addison-Wesley

18.1 White Dwarfs

- Our goals for learning
- What is a white dwarf?
- What can happen to a white dwarf in a close binary system?

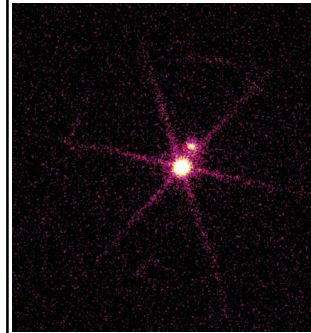
© 2004 Pearson Education, Inc., publishing as Addison-Wesley

What is a white dwarf?



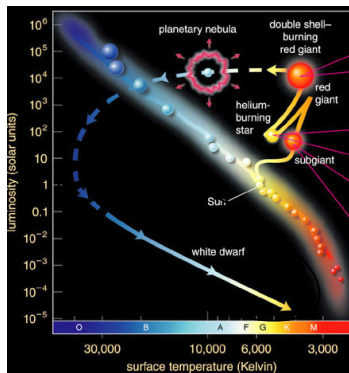
© 2004 Pearson Education, Inc., publishing as Addison-Wesley

White Dwarfs



- White dwarfs are the remaining cores of dead stars
- Electron degeneracy pressure supports them against gravity

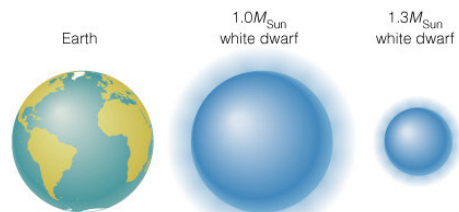
© 2004 Pearson Education, Inc., publishing as Addison-Wesley



White dwarfs cool off and grow dimmer with time

© 2004 Pearson Education, Inc., publishing as Addison-Wesley

Size of a White Dwarf



- White dwarfs with same mass as Sun are about same size as Earth
- Higher mass white dwarfs are smaller

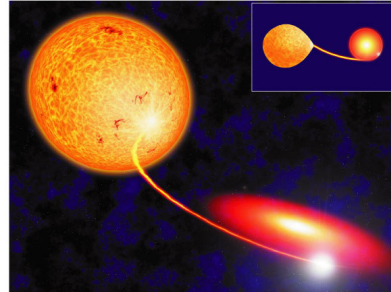
© 2004 Pearson Education, Inc., publishing as Addison-Wesley

The White Dwarf Limit

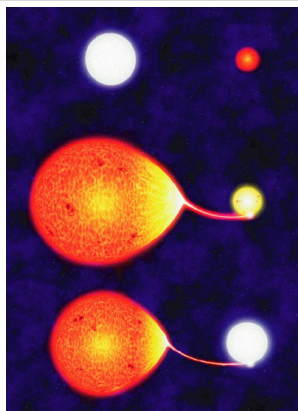
- Quantum mechanics says that electrons must move faster as they are squeezed into a very small space
- As a white dwarf's mass approaches $1.4M_{\text{Sun}}$, its electrons must move at nearly the speed of light
- Because nothing can move faster than light, a white dwarf cannot be more massive than $1.4M_{\text{Sun}}$, the *white dwarf limit* (or *Chandrasekhar limit*)

© 2004 Pearson Education, Inc., publishing as Addison-Wesley

What can happen to a white dwarf in a close binary system?



© 2004 Pearson Education, Inc., publishing as Addison-Wesley



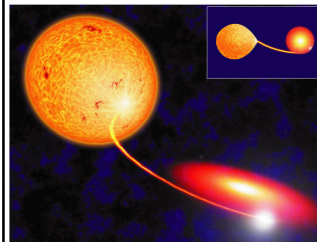
Star that started with less mass gains mass from its companion

Eventually the mass-losing star will become a white dwarf

What happens next?

© 2004 Pearson Education, Inc., publishing as Addison-Wesley

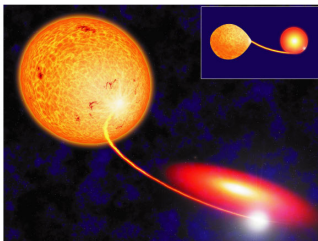
Accretion Disks



- Mass falling toward a white dwarf from its close binary companion has some angular momentum
- The matter therefore orbits the white dwarf in an *accretion disk*

© 2004 Pearson Education, Inc., publishing as Addison-Wesley

Accretion Disks



- Friction between orbiting rings of matter in the disk transfers angular momentum outward and causes the disk to heat up and glow

© 2004 Pearson Education, Inc., publishing as Addison-Wesley

Thought Question

What would gas in disk do if there were no friction?

- It would orbit indefinitely.
- It would eventually fall in.
- It would blow away.

© 2004 Pearson Education, Inc., publishing as Addison-Wesley

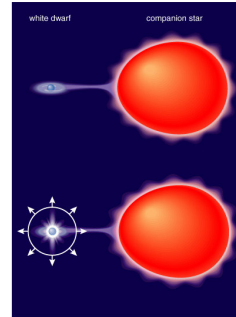
Thought Question

What would gas in disk do if there were no friction?

- A. It would orbit indefinitely.
- B. It would eventually fall in.
- C. It would blow away.

© 2004 Pearson Education, Inc., publishing as Addison-Wesley

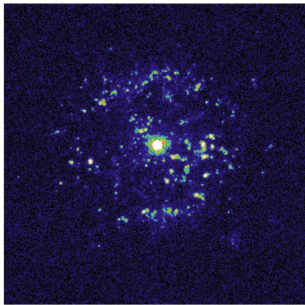
Nova



- The temperature of accreted matter eventually becomes hot enough for hydrogen fusion
- Fusion begins suddenly and explosively, causing a *nova*

© 2004 Pearson Education, Inc., publishing as Addison-Wesley

Nova



- The nova star system temporarily appears much brighter
- The explosion drives accreted matter out into space

© 2004 Pearson Education, Inc., publishing as Addison-Wesley

Thought Question

What happens to a white dwarf when it accretes enough matter to reach the $1.4 M_{\text{Sun}}$ limit?

- A. It explodes
- B. It collapses into a neutron star
- C. It gradually begins fusing carbon in its core

© 2004 Pearson Education, Inc., publishing as Addison-Wesley

Thought Question

What happens to a white dwarf when it accretes enough matter to reach the $1.4 M_{\text{Sun}}$ limit?

- A. It explodes
- B. It collapses into a neutron star
- C. It gradually begins fusing carbon in its core

© 2004 Pearson Education, Inc., publishing as Addison-Wesley

Two Types of Supernova

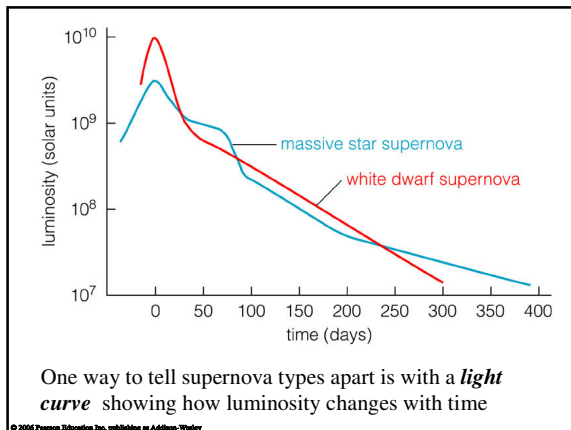
Massive star supernova:

Iron core of massive star reaches white dwarf limit and collapses into a neutron star, causing explosion

White dwarf supernova:

Carbon fusion suddenly begins as white dwarf in close binary system reaches white dwarf limit, causing total explosion

© 2004 Pearson Education, Inc., publishing as Addison-Wesley



Nova or Supernova?

- Supernovae are MUCH MUCH more luminous!!! (about 10 million times)
- Nova: H to He fusion of a layer of accreted matter, white dwarf left intact
- Supernova: complete explosion of white dwarf, nothing left behind

© 2004 Pearson Education, Inc., publishing as Addison-Wesley

Supernova Type: Massive Star or White Dwarf?

- Light curves differ
- Spectra differ (exploding white dwarfs don't have hydrogen absorption lines)

© 2004 Pearson Education, Inc., publishing as Addison-Wesley

What have we learned?

- What is a white dwarf?
 - A white dwarf is the inert core of a dead star
 - Electron degeneracy pressure balances the inward pull of gravity
- What can happen to a white dwarf in a close binary system?
 - Matter from its close binary companion can fall onto the white dwarf through an accretion disk
 - Accretion of matter can lead to novae and white dwarf supernovae

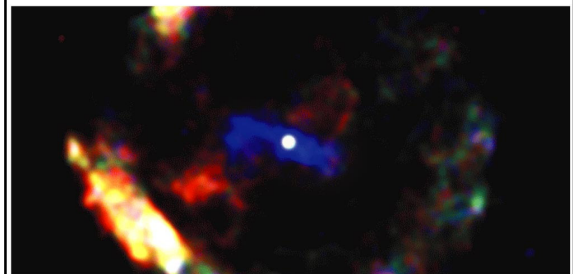
© 2004 Pearson Education, Inc., publishing as Addison-Wesley

18.2 Neutron Stars

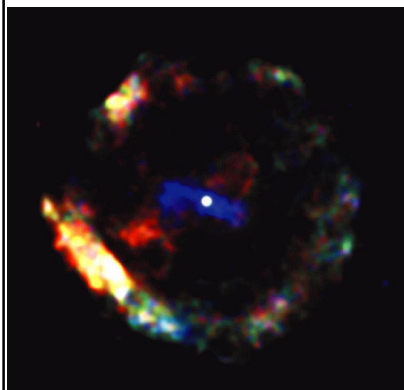
- Our goals for learning
- What is a neutron star?
- How were neutron stars discovered?
- What can happen to a neutron star in a close binary system?

© 2004 Pearson Education, Inc., publishing as Addison-Wesley

What is a neutron star?



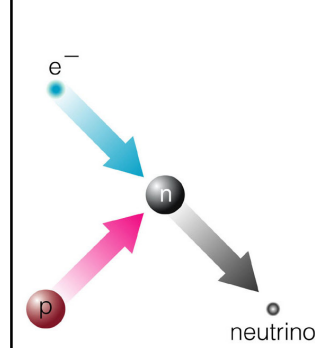
© 2004 Pearson Education, Inc., publishing as Addison-Wesley



A neutron star is the ball of neutrons left behind by a massive-star supernova

Degeneracy pressure of neutrons supports a neutron star against gravity

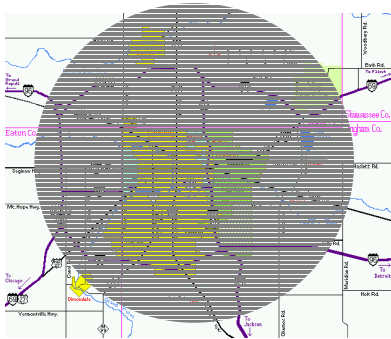
© 2004 Pearson Education, Inc., publishing as Addison-Wesley



Electron degeneracy pressure goes away because electrons combine with protons, making neutrons and neutrinos

Neutrons collapse to the center, forming a **neutron star**

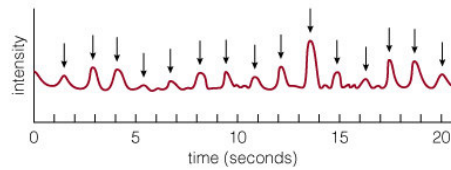
© 2004 Pearson Education, Inc., publishing as Addison-Wesley



A neutron star is about the same size as a small city

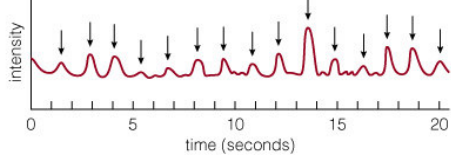
© 2004 Pearson Education, Inc., publishing as Addison-Wesley

How were neutron stars discovered?



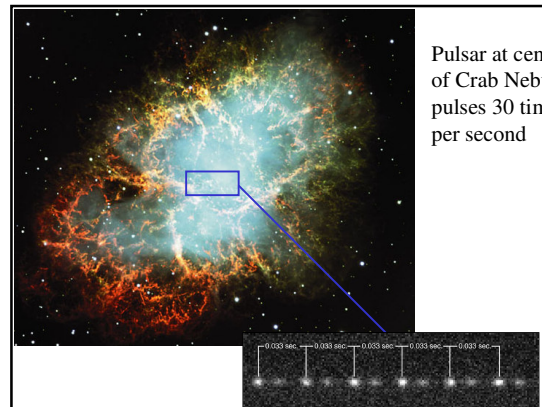
© 2004 Pearson Education, Inc., publishing as Addison-Wesley

Discovery of Neutron Stars



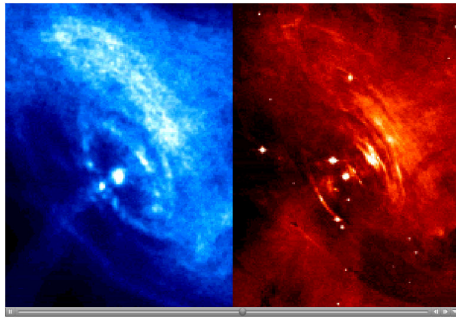
- Using a radio telescope in 1967, Jocelyn Bell noticed very regular pulses of radio emission coming from a single part of the sky
- The pulses were coming from a spinning neutron star—a *pulsar*

© 2004 Pearson Education, Inc., publishing as Addison-Wesley



Pulsar at center of Crab Nebula pulses 30 times per second

© 2004 Pearson Education, Inc., publishing as Addison-Wesley

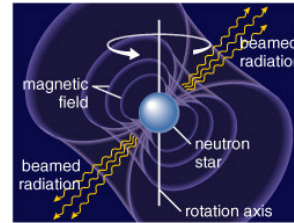


X-rays

Visible light

© 2004 Pearson Education, Inc., publishing as Addison-Wesley

Pulsars



- A pulsar is a neutron star that beams radiation along a magnetic axis that is not aligned with the rotation axis

© 2004 Pearson Education, Inc., publishing as Addison-Wesley

Pulsars



- The radiation beams sweep through space like lighthouse beams as the neutron star rotates

© 2004 Pearson Education, Inc., publishing as Addison-Wesley

Why Pulsars must be Neutron Stars

Circumference of NS = 2π (radius) ~ 60 km

Spin Rate of Fast Pulsars ~ 1000 cycles per second

Surface Rotation Velocity $\sim 60,000$ km/s
 $\sim 20\%$ speed of light
 \sim escape velocity from NS

Anything else would be torn to pieces!

© 2004 Pearson Education, Inc., publishing as Addison-Wesley

Collapse of the Solar Nebula



Pulsars spin fast because core's spin speeds up as it collapses into neutron star

Conservation of angular momentum

© 2004 Pearson Education, Inc., publishing as Addison-Wesley

Thought Question

Could there be neutron stars that appear as pulsars to other civilizations but not to us?

- A. Yes
- B. No

© 2004 Pearson Education, Inc., publishing as Addison-Wesley

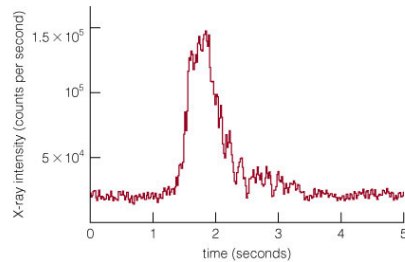
Thought Question

Could there be neutron stars that appear as pulsars to other civilizations but not to us?

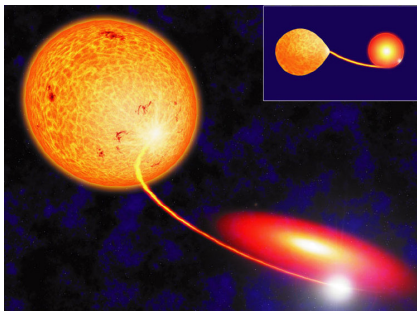
- A. Yes
- B. No

© 2004 Pearson Education, Inc., publishing as Addison-Wesley

What can happen to a neutron star in a close binary system?

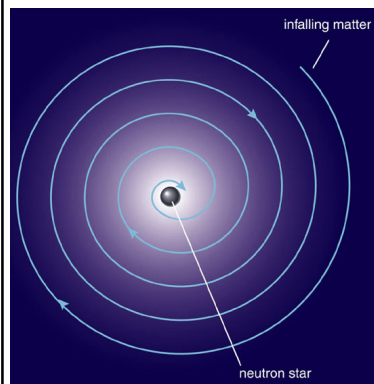


© 2004 Pearson Education, Inc., publishing as Addison-Wesley



Matter falling toward a neutron star forms an accretion disk, just as in a white-dwarf binary

© 2004 Pearson Education, Inc., publishing as Addison-Wesley



Accreting matter adds angular momentum to a neutron star, increasing its spin

Episodes of fusion on the surface lead to X-ray bursts

© 2004 Pearson Education, Inc., publishing as Addison-Wesley

Thought Question

According to conservation of angular momentum, what would happen if a star orbiting in a direction opposite the neutron's star rotation fell onto a neutron star?

- A. The neutron star's rotation would speed up.
- B. The neutron star's rotation would slow down.
- C. Nothing, the directions would cancel each other out.

© 2004 Pearson Education, Inc., publishing as Addison-Wesley

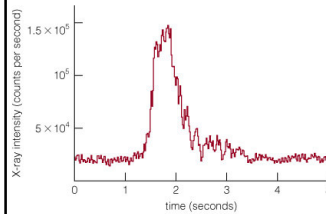
Thought Question

According to conservation of angular momentum, what would happen if a star orbiting in a direction opposite the neutron's star rotation fell onto a neutron star?

- A. The neutron star's rotation would speed up.
- B. The neutron star's rotation would slow down.**
- C. Nothing, the directions would cancel each other out.

© 2004 Pearson Education, Inc., publishing as Addison-Wesley

X-Ray Bursts



- Matter accreting onto a neutron star can eventually become hot enough for helium fusion
- The sudden onset of fusion produces a burst of X-rays

© 2004 Pearson Education, Inc., publishing as Addison-Wesley

What have we learned?

- What is a neutron star?
 - A ball of neutrons left over from a massive star supernova and supported by neutron degeneracy pressure
- How were neutron stars discovered?
 - Beams of radiation from a rotating neutron star sweep through space like lighthouse beams, making them appear to pulse
 - Observations of these pulses were the first evidence for neutron stars

© 2004 Pearson Education, Inc., publishing as Addison-Wesley

What have we learned?

- What can happen to a neutron star in a close binary system?
 - The accretion disk around a neutron star gets hot enough to produce X-rays, making the system an X-ray binary
 - Sudden fusion events periodically occur on a the surface of an accreting neutron star, producing X-ray bursts

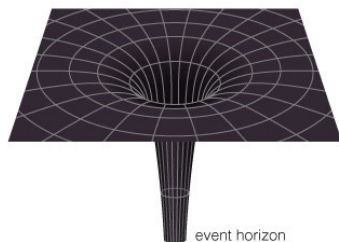
© 2004 Pearson Education, Inc., publishing as Addison-Wesley

18.3 Black Holes: Gravity's Ultimate Victory

- Our goals for learning
- What is a black hole?
- What would it be like to visit a black hole?
- Do black holes really exist?

© 2004 Pearson Education, Inc., publishing as Addison-Wesley

What is a black hole?



© 2004 Pearson Education, Inc., publishing as Addison-Wesley

A **black hole** is an object whose gravity is so powerful that not even light can escape it.

© 2004 Pearson Education, Inc., publishing as Addison-Wesley

Thought Question

What happens to the escape velocity from an object if you shrink it?

- A. It increases
- B. It decreases
- C. It stays the same

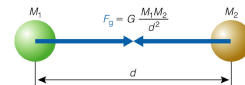
© 2004 Pearson Education, Inc., publishing as Addison-Wesley

Thought Question

What happens to the escape velocity from an object if you shrink it?

- A. It increases
- B. It decreases
- C. It stays the same

Hint:



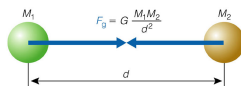
© 2004 Pearson Education, Inc., publishing as Addison-Wesley

Thought Question

What happens to the escape velocity from an object if you shrink it?

- A. It increases**
- B. It decreases
- C. It stays the same

Hint:



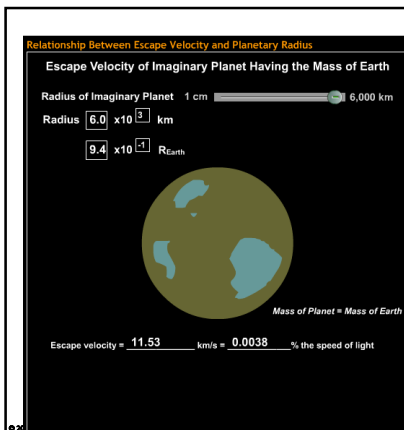
© 2004 Pearson Education, Inc., publishing as Addison-Wesley

Escape Velocity

Initial Kinetic Energy = Final Gravitational Potential Energy

$$\frac{(\text{escape velocity})^2}{2} = \frac{G \times (\text{mass})}{(\text{radius})}$$

© 2004 Pearson Education, Inc., publishing as Addison-Wesley



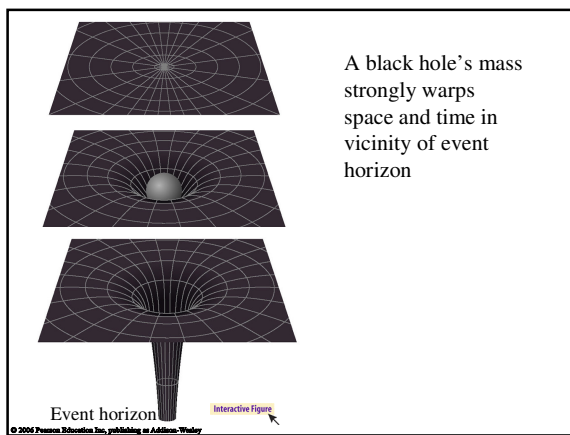
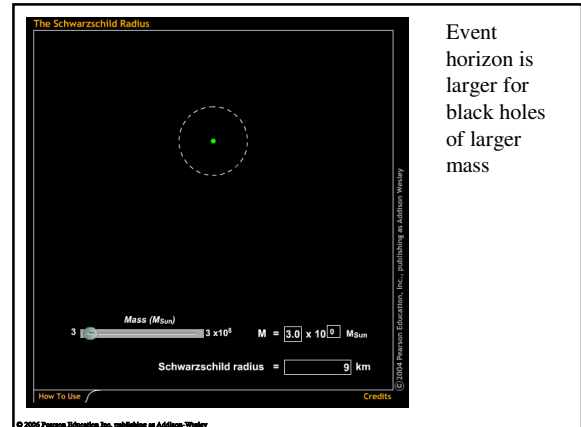
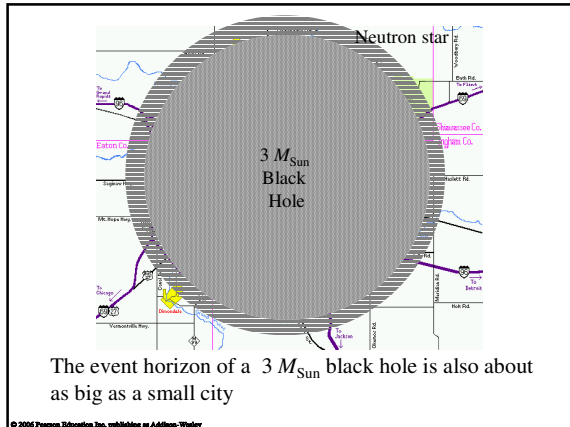
Light would not be able to escape Earth's surface if you could shrink it to < 1 cm

© 2004 Pearson Education, Inc., publishing as Addison-Wesley

“Surface” of a Black Hole

- The “surface” of a black hole is the radius at which the escape velocity equals the speed of light.
- This spherical surface is known as the *event horizon*.
- The radius of the event horizon is known as the *Schwarzschild radius*.

© 2004 Pearson Education, Inc., publishing as Addison-Wesley



No Escape

- Nothing can escape from within the event horizon because nothing can go faster than light.
- No escape means there is no more contact with something that falls in. It increases the hole mass, changes the spin or charge, but otherwise loses its identity.

© 2004 Pearson Education, Inc., publishing as Addison-Wesley

Neutron Star Limit

- Quantum mechanics says that neutrons in the same place cannot be in the same state
- Neutron degeneracy pressure can no longer support a neutron star against gravity if its mass exceeds about $3 M_{\text{sun}}$
- Some massive star supernovae can make black hole if enough mass falls onto core

© 2004 Pearson Education, Inc., publishing as Addison-Wesley

Singularity

- Beyond the neutron star limit, no known force can resist the crush of gravity.
- As far as we know, gravity crushes all the matter into a single point known as a *singularity*.

© 2004 Pearson Education, Inc., publishing as Addison-Wesley

Thought Question

How does the radius of the event horizon change when you add mass to a black hole?

- A. Increases
- B. Decreases
- C. Stays the same

© 2004 Pearson Education Inc., publishing as Addison-Wesley

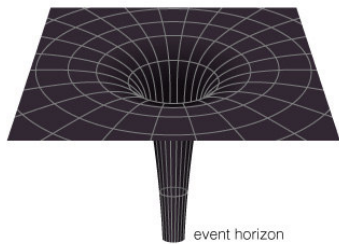
Thought Question

How does the radius of the event horizon change when you add mass to a black hole?

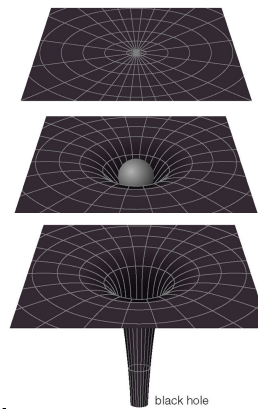
- A. Increases
- B. Decreases
- C. Stays the same

© 2004 Pearson Education Inc., publishing as Addison-Wesley

What would it be like to visit a black hole?



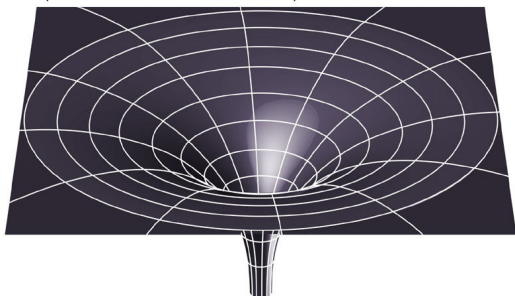
© 2004 Pearson Education Inc., publishing as Addison-Wesley



If the Sun shrank into a black hole, its gravity would be different only near the event horizon

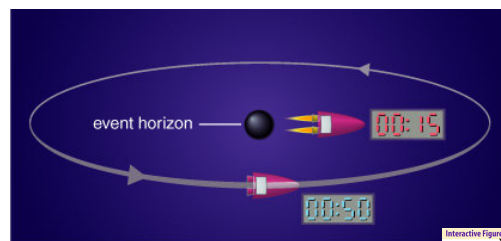
Black holes don't suck!

© 2004 Pearson Education Inc., publishing as Addison-Wesley



Light waves take extra time to climb out of a deep hole in spacetime leading to a ***gravitational redshift***

© 2004 Pearson Education Inc., publishing as Addison-Wesley



Time passes more slowly near the event horizon

© 2004 Pearson Education Inc., publishing as Addison-Wesley

Thought Question

Is it easy or hard to fall into a black hole?

- A. Easy
- B. Hard

© 2006 Pearson Education, Inc., publishing as Addison-Wesley

Thought Question

Is it easy or hard to fall into a black hole?

- A. Easy
- B. Hard

Hint: A black hole with the same mass as the Sun wouldn't be much bigger than a college campus

© 2006 Pearson Education, Inc., publishing as Addison-Wesley

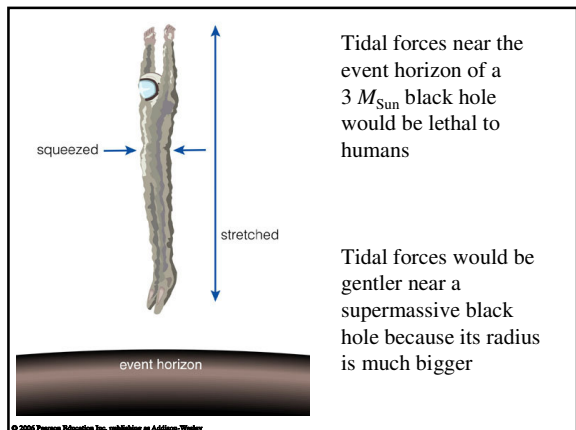
Thought Question

Is it easy or hard to fall into a black hole?

- B. Hard

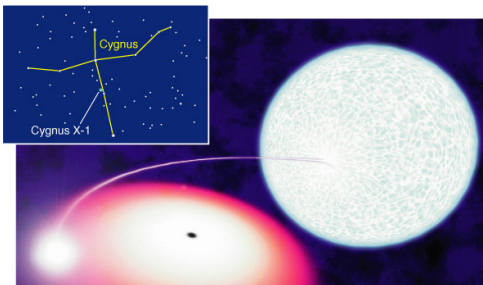
Hint: A black hole with the same mass as the Sun wouldn't be much bigger than a college campus

© 2006 Pearson Education, Inc., publishing as Addison-Wesley



© 2006 Pearson Education, Inc., publishing as Addison-Wesley

Do black holes really exist?

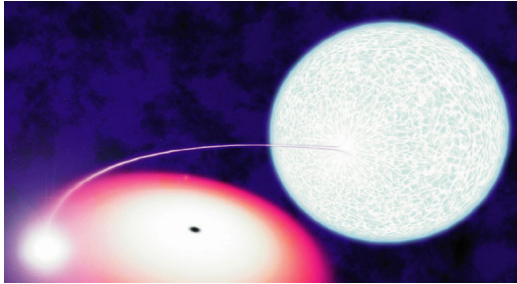


© 2006 Pearson Education, Inc., publishing as Addison-Wesley

Black Hole Verification

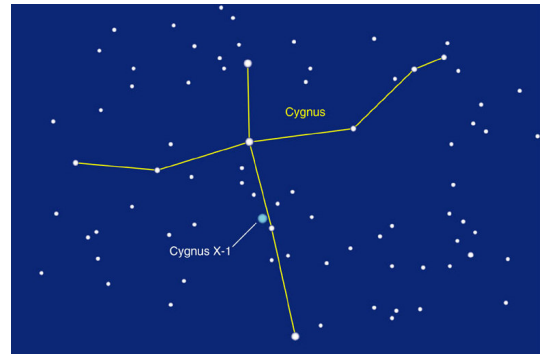
- Need to measure mass
 - Use orbital properties of companion
 - Measure velocity and distance of orbiting gas
- It's a black hole if it's not a star and its mass exceeds the neutron star limit ($\sim 3 M_{\text{Sun}}$)

© 2006 Pearson Education, Inc., publishing as Addison-Wesley



Some X-ray binaries contain compact objects of mass exceeding $3 M_{\text{Sun}}$ which are likely to be black holes

© 2006 Pearson Education, Inc., publishing as Addison-Wesley



One famous X-ray binary with a likely black hole is in the constellation Cygnus

© 2006 Pearson Education, Inc., publishing as Addison-Wesley

What have we learned?

- What is a black hole?
 - A black hole is a massive object whose radius is so small that the escape velocity exceeds the speed of light
- What would it be like to visit a black hole?
 - You can orbit a black hole like any other object of the same mass—black holes don't suck!
 - Near the event horizon time slows down and tidal forces are very strong

© 2006 Pearson Education, Inc., publishing as Addison-Wesley

What have we learned?

- Do black holes really exist?
 - Some X-ray binaries contain compact objects so massive to be neutron stars—they are almost certainly black holes

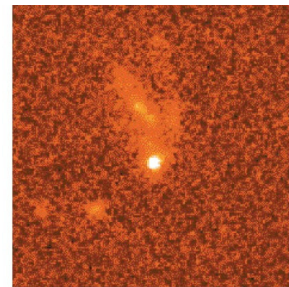
© 2006 Pearson Education, Inc., publishing as Addison-Wesley

18.4 The Mystery of Gamma Ray Bursts

- Our goals for learning
- Where do gamma-ray bursts come from?
- What causes gamma-ray bursts?

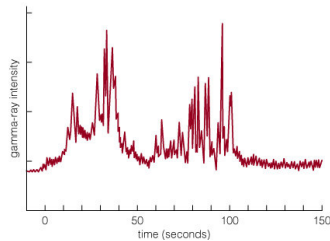
© 2006 Pearson Education, Inc., publishing as Addison-Wesley

Where do gamma-ray bursts come from?



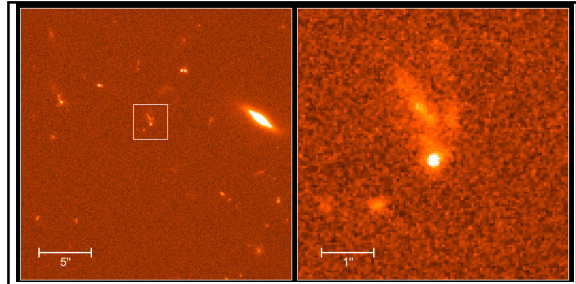
© 2006 Pearson Education, Inc., publishing as Addison-Wesley

Gamma-Ray Bursts



- Brief bursts of gamma-rays coming from space were first detected in the 1960s

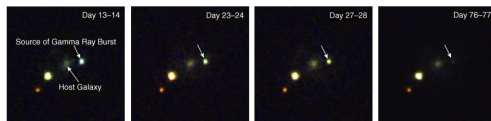
© 2004 Pearson Education, Inc., publishing as Addison-Wesley



- Observations in the 1990s showed that many gamma-ray bursts were coming from very distant galaxies
- They must be among the most powerful explosions in the universe—could be the formation of a black hole

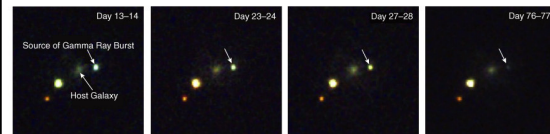
© 2004 Pearson Education, Inc., publishing as Addison-Wesley

What causes gamma-ray bursts?



© 2004 Pearson Education, Inc., publishing as Addison-Wesley

Supernovae and Gamma-Ray Bursts



- Observations show that at least some gamma-ray bursts are produced by supernova explosions
- Some others may come from collisions between neutron stars

© 2004 Pearson Education, Inc., publishing as Addison-Wesley

What have we learned?

- Where do gamma-ray bursts come from?
 - Most gamma-ray bursts come from distant galaxies
 - They must be among the most powerful explosions in the universe, probably signifying the formation of black holes
- What causes gamma-ray bursts?
 - At least some gamma-ray bursts come from supernova explosions

© 2004 Pearson Education, Inc., publishing as Addison-Wesley